HALF-SPACE ALBEDO PROBLEM IN A RAYLEIGH SCATTERING ATMOSPHERE WITH TRUE ABSORPTION

Menekşe Şenyiğit and Ayşe Kaşkaş Ankara University Faculty of Science Department of Physics, 06100 Tandoğan Ankara, Turkey

ABSTRACT

The H_N method, employed for studies in neutron transport theory, is used to establish numerical results basic to the vector equation describing the transfer of polarized light in a Rayleigh scattering atmosphere with true absorption^{1,2}. The method has been applied to the half-space albedo problem.

We consider the vector equation of transfer in plane geometry^{3,4,}

$$\mu \frac{\partial \mathbf{I}(\tau,\mu)}{\partial \tau} + \mathbf{I}(\tau,\mu) = \frac{1}{2} \omega \mathbf{Q}(\mu) \int_{-1}^{1} \mathbf{Q}^{\mathrm{T}}(\mu') \mathbf{I}(\tau,\mu') d\mu'$$
(1)

 $I(\tau,\mu)$ is a vector whose two components $I_{\ell}(\tau,\mu)$ and $I_{r}(\tau,\mu)$ are the angular intensities in the two states of polarization. τ is the optical variable, μ is the direction cosine (as measured from the positive τ axis) of the propagating radiation, and $\omega \in [0,1]$ is the singlescattering albedo. Here $Q^{T}(\mu)$ denotes the transpose of $Q(\mu)^{5}$

$$\mathbf{Q}(\mu) = \frac{3}{2}(c+2)^{-1/2} \begin{bmatrix} c\mu^2 + \frac{2}{3}(1-c) & (2c)^{1/2}(1-\mu^2) \\ \frac{1}{3}(c+2) & 0 \end{bmatrix}.$$
(2)

The parameter $c, c \in [0,1]$ is a measure of the Rayleigh component of scattering law: c = 1and $\omega = 1$ would yield Chandrasekhar's conservative Rayleigh-scattering case, $\omega = 1$ and $c \in [0,1]$ is Chandrasekhar's conservative case for a mixture of Rayleigh and isotropic scattering laws. c = 1 and $\omega \in [0,1]$ yields general Rayleigh scattering⁶. The general solution of Eq.(1) can be written as

$$\mathbf{I}(\tau,\mu) = A(\eta_0) \mathbf{\Phi}(\eta_0,\mu) e^{-\frac{\tau}{\eta_0}} + A(-\eta_0) \mathbf{\Phi}(-\eta_0,\mu) e^{\frac{\tau}{\eta_0}} + \int_{-1}^{1} \left[A_1(\eta) \mathbf{\Phi}_1(\eta,\mu) + A_2(\eta) \mathbf{\Phi}_2(\eta,\mu) \right] e^{-\frac{\tau}{\eta}} d\eta$$
(3)

here $A(\pm \eta_0)$ and the vector $A_1(\eta), A_2(\eta)$ are the arbitrary expansion coefficients to be determined once the appropriate boundary conditions are specified⁷. The discrete and continuum normal modes are written respectively as

$$\mathbf{\Phi}(\pm\eta_0,\mu) = \frac{1}{2}\omega\eta_0 \frac{1}{(\eta_0 \mp \mu)} \mathbf{Q}(\mu) \mathbf{M}(\pm\eta_0)$$
(4a)

$$\boldsymbol{\Phi}_{\alpha}(\eta,\mu) = \frac{1}{2}\omega \left[\eta \frac{P}{\eta-\mu} + \lambda_{\alpha}(\eta)\delta(\eta-\mu)\right] \mathbf{Q}(\mu) \mathbf{M}_{\alpha}(\eta)$$
(4b)

Let us consider the half-space medium at x > 0 and vacuum at x < 0. In Eq.(1) neglecting $A(-\eta_0)$ and $A_1(-\eta)$, $A_2(-\eta)$, we write the required solution as

$$\mathbf{I}(0,\mu) = A(\eta_0) \mathbf{\Phi}(\eta_0,\mu) + \int_0^1 \left[A_1(\eta) \mathbf{\Phi}_1(\eta,\mu) + A_2(\eta) \mathbf{\Phi}_2(\eta,\mu) \right] d\eta \qquad .$$
(5)

For the half-space albedo problem we seek the solution of Eq.(5) which satisfies the following boundary condition⁸

$$\mathbf{I}(0,\mu) = \delta(\mu - \mu_0)\mathbf{F} \qquad , \quad \mu, \, \mu_0 > 0 \tag{6}$$

where \mathbf{F} is a constant vector. The expansion of exit distribution can be chosen as ⁹

$$\mathbf{I}(0,-\mu) = \mathbf{Q}(\mu) \sum_{\alpha=0}^{N} \mathbf{a}_{\alpha} \mu^{\alpha} \quad , \quad \mu > 0$$
⁽⁷⁾

To get the expansion coefficients in Eq.(5), it is multiplied by $\mu \Phi^{T}(\xi,\mu)$, $(\xi = \eta_0, \pm \eta)$, integrated over μ , $\mu \in (-1,+1)$ and used the orthogonality relations⁵

$$A(\eta_0) = -\frac{1}{N(\eta_0)} \left[\frac{\omega \eta_0}{2} \sum_{\alpha=0}^N \mathbf{\Gamma}_{\alpha}(\eta_0) \mathbf{a}_{\alpha} + \mu_0 \mathbf{\Phi}^T(\eta_0, \mu_0) \mathbf{F} \right]$$
(8)

$$A_{1}(\eta) = \frac{1}{N(\eta)} \left[-\frac{\omega\eta}{2} \sum_{\alpha=0}^{N} \left(N_{22}(\eta) \Gamma_{\alpha}^{(1)}(\eta) - N_{12}(\eta) \Gamma_{\alpha}^{(2)}(\eta) \right) \mathbf{a}_{\alpha} + \mu_{0} \mathbf{\Phi}_{1}^{T^{\dagger}}(\eta, \mu_{0}) \mathbf{F} \right]$$
(9)

$$A_{2}(\eta) = \frac{1}{N(\eta)} \left[-\frac{\omega\eta}{2} \sum_{\alpha=0}^{N} \left(-N_{21}(\eta) \Gamma_{\alpha}^{(1)}(\eta) + N_{11}(\eta) \Gamma_{\alpha}^{(2)}(\eta) \right) \mathbf{a}_{\alpha} + \mu_{0} \Phi_{2}^{T^{\dagger}}(\eta, \mu_{0}) \mathbf{F} \right]$$
(10)

here

$$\Gamma_{\alpha}(\eta_0) = \frac{2}{\omega\eta_0} \int_0^1 \mu^{\alpha+1} \mathbf{\Phi}^{\mathrm{T}}(\eta_0, -\mu) \mathbf{Q}(\mu) d\mu$$
(11a)

$$\Gamma^{\beta}_{\alpha}(\eta) = \frac{2}{\omega\eta} \int_{0}^{1} \mu^{\alpha+1} \Phi_{\beta}^{T}(\eta, -\mu) \mathbf{Q}(\mu) d\mu, \qquad \beta = 1, 2.$$
(11b)

Now we consider the exit distribution in Eq.(5) replacing μ with $-\mu$ and multiplying it by μ^{m+1} , integrating over $\mu \in (0,1)$ we obtain

$$\sum_{\alpha=0}^{N} \left\{ \int_{0}^{1} \mu^{m+\alpha+1} \mathbf{Q}(\mu) d\mu + \left(\frac{\omega\eta_{0}}{2}\right)^{2} \frac{1}{N(\eta_{0})} \mathbf{A}_{m}(\eta_{0}) \mathbf{\Gamma}_{\alpha}(\eta_{0}) - \left(\frac{\omega}{\eta}\right)^{2} \int_{0}^{1} \frac{\eta^{2}}{N(\eta)} \mathbf{A}_{m}^{1}(\eta) [-N_{22}(\eta) \mathbf{\Gamma}_{\alpha}^{1}(\eta) + N_{12}(\eta) \mathbf{\Gamma}_{\alpha}^{2}(\eta)] d\eta - \left(\frac{\omega}{\eta}\right)^{2} \int_{0}^{1} \frac{\eta^{2}}{N(\eta)} \mathbf{A}_{m}^{2}(\eta) [N_{21}(\eta) \mathbf{\Gamma}_{\alpha}^{1}(\eta) - N_{11}(\eta) \mathbf{\Gamma}_{\alpha}^{2}(\eta)] d\eta \right\} \mathbf{a}_{\alpha}$$

$$= \frac{\omega\eta_{0}}{2} \mathbf{A}_{m}(\eta_{0}) \frac{1}{N(\eta_{0})} \mu_{0} \mathbf{\Phi}^{T}(\eta_{0}, \mu_{0}) \mathbf{F} + \int_{0}^{1} \frac{\omega\eta}{2} \frac{1}{N(\eta)} \mu_{0} \left(\mathbf{A}_{m}^{(1)}(\eta) \mathbf{\Phi}_{1}^{T^{\dagger}}(\eta, \mu) + \mathbf{A}_{m}^{(2)}(\eta) \mathbf{\Phi}_{2}^{T^{\dagger}}(\eta, \mu)\right) \mathbf{F} d\eta$$

$$(12)$$

$$\mathbf{A}_{\alpha}(\eta_{0}) = \frac{2}{\omega\eta_{0}} \int_{0}^{1} \mu^{\alpha+1} \mathbf{\Phi}(-\eta_{0},\mu) d\mu$$

$$\mathbf{A}_{\alpha}^{\beta}(\eta) = \frac{2}{\omega\eta} \int_{0}^{1} \mu^{\alpha+1} \mathbf{\Phi}_{\alpha}^{\beta}(-\eta_{0},\mu) d\mu \quad , \quad \beta = 1,2$$
(13)

The numerical results of the required albedo values, as functions of the single scattering albedo and c which is a measure of the Rayleigh component of scattering law, can be obtained from the following equation

$$\beta = \int_{0}^{1} \begin{bmatrix} 1 \\ 1 \end{bmatrix}^{T} \mathbf{I}(0, -\mu) \, \mu \, d\mu \, \bigg/ \int_{0}^{1} \begin{bmatrix} 1 \\ 1 \end{bmatrix}^{T} \mathbf{I}(0, \mu) \, \mu \, d\mu \tag{14}$$

where \mathbf{a}_{q} can be found from Eq. (12).

REFERENCES

- 1. Tezcan, C. Kaşkaş, A. and Güleçyüz, M. Ç., The H_N Method for Solving Linear Transport Equation: Theory and Applications, *J. Quant. Spectrosc. Radiat. Transfer*, Vol. 78, pp 243-254, 2003.
- 2. Karahasanoğlu Şenyiğit, M. and Kaşkaş, A., The H_N Method for Milne Problem with Polarization, *Astrophysics and Space Science*, (In press).
- 3. Mourad, S., A. and Siewert, C., E., On Solutions of An Equation of Transfer for A Planetary Atmosphere, *Astrophys. J*, Vol. 155, pp 555-564, 1969.
- 4. Schnatz, T., W. and Siewert, C., E., Radiative Transfer in A Rayleigh-Scattering Atmosphere with True Absorption, *J. Math.Phys,* Vol. 11, pp 2733-2739, 1970.
- Bond, G. R. and Siewert, C. E., On the Nonconservative Equation of Transfer for A Combination of Rayleigh and Isotropic Scattering, *Astrophys. J*, Vol. 164, pp 97-110, 1971.
- 6. Siewert, C. E., A Concise and Accurate Solution for A Polarization Model in Radiative Transfer, *J. Quant. Spectrosc. Radiat. Transfer*, Vol. 62, pp 677-684, 1999.
- 7. Burniston, E. E. and Siewert, C., E., Half-Range Expansion Theorems in Studies of Polarized Light, *J. Quant. Spectrosc. Radiat. Transfer*, Vol. 11, pp 3416-3420, 1970.
- 8. Schnatz, T. W. and Siewert, C., E., On the Transfer of Polarized Light in Rayleigh-Scattering Half Spaces with True Absorption, *Mon. Not. R. Astr. Soc.*, Vol. 152, pp 491-508, 1971.
- Siewert, C. E., On Using the F_N Method for Polarization Studies in Finite Planeparallel Atmospheres, *J. Quant. Spectrosc. Radiat. Transfer*, Vol. 21, pp 35-39, (1979).